Planetary boundary layer

Lecture 2

Homework

- 1-5 Cátia
- 2-2 Diogo
- 2-3 Florian
- 2-5 Jason
- 2-8 Maria
- 2-9 Mariana
- 2-11 Sara

The equations in a continuum (10 eq, 10 unknows)

$$\begin{split} \frac{\partial \vec{v}}{\partial t} &= -\vec{v}. \nabla \vec{v} - \frac{1}{\rho} \nabla P + \vec{g} - 2\vec{\Omega} \times \vec{v} + \nu \nabla^2 \vec{v} & \text{Navier-Stokes} \\ \frac{\partial \theta}{\partial t} &= -\vec{v}. \nabla \theta + \dot{Q}_{rad} + \dot{Q}_{lat} + \kappa \nabla^2 \theta & \text{Thermodynamics} \\ \frac{\partial \rho}{\partial t} &= -\nabla. \left(\rho \vec{v}\right) = -\vec{v}. \nabla \rho - \rho \nabla. \vec{v} & \text{Continuity} \\ \frac{\partial q_{v,l,s,\dots}}{\partial t} &= -\vec{v}. \nabla q_{v,l,s,\dots} + G_{q_{v,l,s,\dots}} + \kappa_D \nabla^2 q_{v,l,s,\dots} & \text{Water budget} \\ p &= R_d \rho T (1 + 0.61 q_v) & \text{State equation} \\ \theta &= T \left(\frac{P}{P_{00}}\right)^{-\frac{R_d}{c_p}} & \text{Potential temperature} \end{split}$$

The Navier-Stokes equation is an application of Newton's second law

$$\vec{F} = m\vec{a} \Longrightarrow \vec{a} = \frac{d\vec{v}}{dt} = \frac{\vec{F}}{m}$$

it applies to a moving body of mass m. In a fluid we can not follow a set of fluid molecules. Euler proposed to look instead at fixed (not moving) elements of volume. The time evolution of a variable at a fixed position is its tendency, given by the partial derivative $\partial/\partial t$. Euler showed that:

$$\frac{d\vec{v}}{dt} = \frac{\partial\vec{v}}{\partial t} + u\frac{\partial\vec{v}}{\partial t} + v\frac{\partial\vec{v}}{\partial y} + w\frac{\partial\vec{v}}{\partial z}$$

Where $\vec{v} \equiv (u, v, w)$.

This leads to the Euler equation:

$$-\vec{v} \cdot \nabla \vec{v}$$
$$\frac{\partial \vec{v}}{\partial t} = -u \frac{\partial \vec{v}}{\partial t} - v \frac{\partial \vec{v}}{\partial y} - w \frac{\partial \vec{v}}{\partial z} + \frac{\vec{F}}{m}$$
tendency advection

$$\frac{\partial \vec{v}}{\partial t} = -\vec{v} \cdot \nabla \vec{v} - \frac{1}{\rho} \nabla P + \vec{g} - 2\vec{\Omega} \times \vec{v} + \nu \nabla^2 \vec{v}$$

Navier and Stokes introduced viscosity. Viscosity changes the order of the equation and allows the no-slip condition at the boundary.

Within the fluid, viscosity promotes diffusion: the transport of properties leading to homogeneization. It appears in different equations:

$$\mathbf{v}\nabla^2 \vec{v}; \mathbf{\kappa}\nabla^2 \theta; \mathbf{\kappa}_D \nabla^2 q_{v,l,s,\dots}$$

If a given variable is spatially homogenous, $\nabla^2 \theta = 0$ and this terms are zero.

The equations can only be solved numerically

The equations are non-linear. In the Navier-Stokes equation, non-linearity comes from 2 terms:

$$\frac{\partial \vec{v}}{\partial t} = -\vec{v} \cdot \nabla \vec{v} - \frac{1}{\rho} \nabla P + \vec{g} - 2\vec{\Omega} \times \vec{v} + \nu \nabla^2 \vec{v}$$

To solve the equations (by analytical or numerical methods) we need initial and boundary conditions: we need observations of the initial state, and other equations for the Earth and Ocean surface. We cannot observe the continuum

Reynolds noted that to relate the equations with real world states, one needs to apply spatial/time averaging: that is what sensors do.

Reynolds averaging

Equations

Approximations by scaling: compare the relevance of different terms

Perturbation methods: u = U + u'Linearization: compute $\frac{\partial u'}{\partial t}$ (with $u' \ll U$) Turbulence: compute $\frac{\partial \overline{u}}{\partial t}$