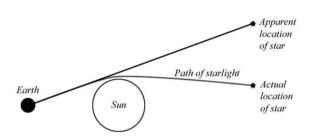
# **Overview of the cosmological model**

**Summary** 

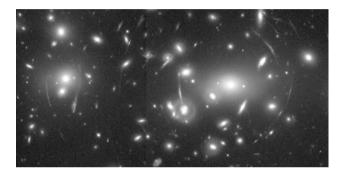
# Physical Cosmology: a physical model for the Universe

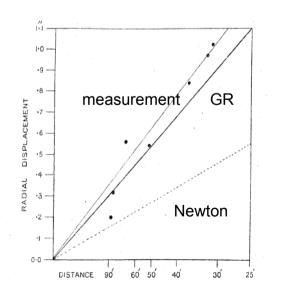
# Gravity (General Relativity)

### Tested on various scales

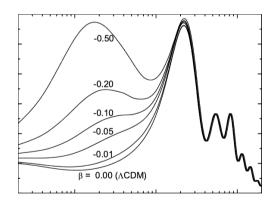








- $\rightarrow$  metric
- $\rightarrow$  Einstein equations

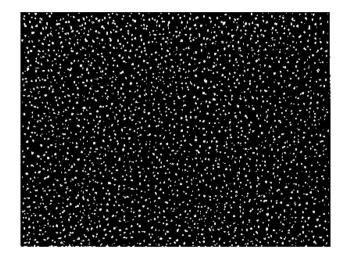


Awarded to **Albert Einstein** "for his services to Theoretical Physics, (and especially for his discovery of the law of the photoelectric effect)." (1/1)

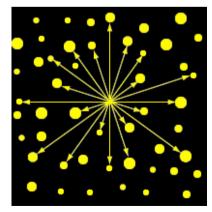


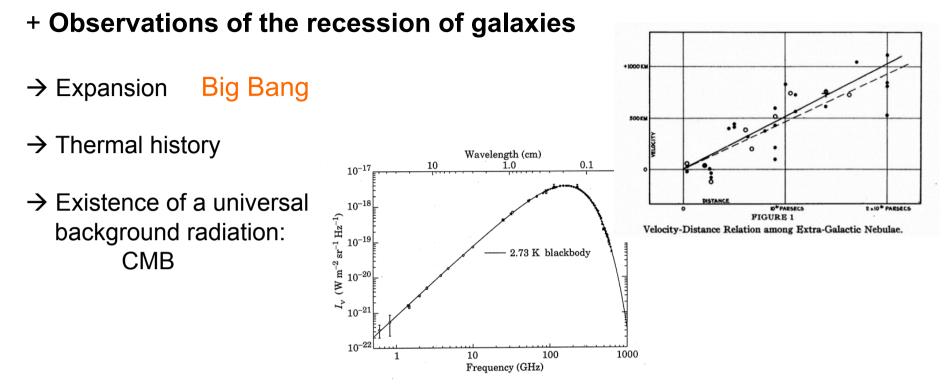
# + Cosmological principle

→ metric is Robertson-Walker, spherically symmetric with two degrees of freedom: a(t), K → and two related cosmological parameters: H<sub>0</sub>,  $\Omega_{K}$ 

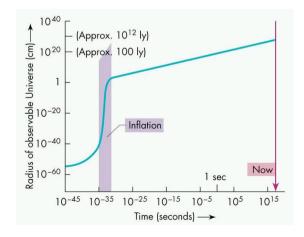


- + Olbers paradox
  - $\rightarrow$  a(t) must vary





 $\rightarrow$  horizon, flatness and coincidence problems: solved by the **Inflation** mechanism



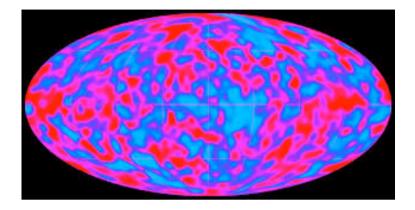
Awarded to Arno A. Penzias and Robert W. Wilson "for their discovery of cosmic microwave background radiation." (1/4 + 1/4)



# + Observation of anisotropies in the CMB

→ Existence of perturbations to the cosmological principle

→ Problem of the origin of the perturbations



Solved by the mechanism of quantum fluctuations + inflation + + gravitational collapse

> Awarded to John C. Mather and George F. Smoot "for their discovery of the blackbody form and anisotropy of the cosmic microwave background radiation." (1/2 + 1/2)

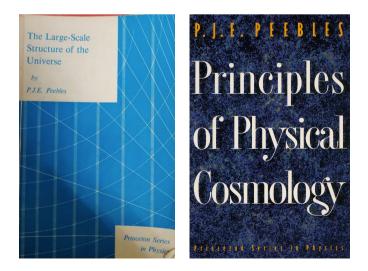


## + Measurement of the anisotropies in the CMB

Their amplitude is very small  $\delta_T \sim 10^{-5} \rightarrow \text{ very small clustering at } z=1100 \ (\delta_b \sim 10^{-5})$ 

- + Gravitational collapse is small ( $\delta_b$  grows only a factor ~10<sup>3</sup> until z=0)
- + There are structures with large density contrast  $\delta$  (large clustering at z=0)
  - $\rightarrow$  Problem of the mechanism of structure formation

Solved by the hypothesis of the existence of an extra component in the cosmological fluid - Dark matter  $\rightarrow$  CDM model



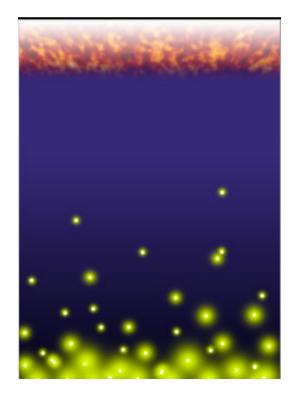
Awarded to **P. James. E. Peebles** "for theoretical discoveries in physical cosmology" (1/2)



### + Non-linear gravitational collapse

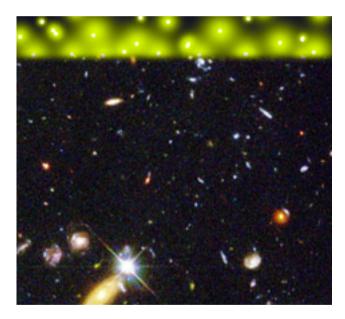
 $\rightarrow$  Formation of dark matter halos and collapse of baryonic matter on those halos (neutral Hydrogen HI clouds that condense and form the first stars)

→ New radiation ionizes the HI clouds, forming inonized Hydrogen regions HII - the reionization of the Universe



## + Galaxy formation

 $\rightarrow$  The gravitational collapse does not describe all aspects of structure formation. Non-gravitational effects associated to the baryonic matter start to be important at this stage:



Cooling - the has gas to cool-down to condense, losing pressure falls into the center of the halo where it can form stars. Angular momentum conservation during the fall produces a disk → spiral galaxies

Feedback - the quantity of cold gas available decreases by influence of the environment

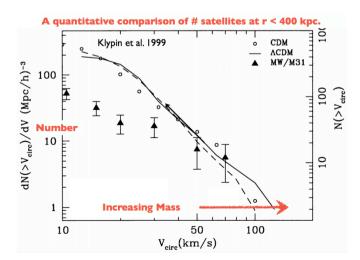
Mergers - frequent interactions between halos may form elliptical galaxies from primitive spiral galaxies.

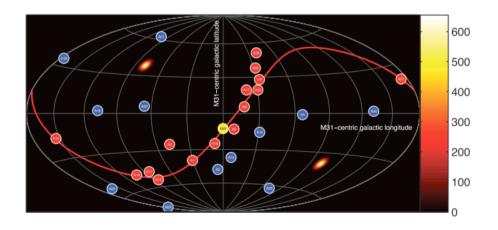
### + Observations of properties of small-scale structure (kpc)

 $\rightarrow$  Problem of the radial density profile of structures (cusp/core)

- → Problem of lack of structures (satellite galaxies)
- → Problem of the satellite orbital plane

Not yet solved - hypothesis of existence of other types of dark matter (Warm Dark Matter, Interacting DM), interacting DM/baryons in dense environments (Baryon feedback), hypothesis of modifications to GR on 'small scales' (MOND)





### + Measurements of distances to Supernovas

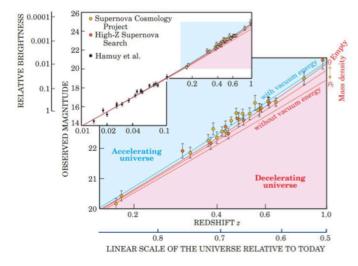
SN at all redshifts are fainter (more distant?) than expected from the  $d_{L}(z)$  predicted by the CDM cosmological model

 $\rightarrow$  The Universe changed from a decelerated expansion to an accelerated one

Solved by assuming the existence of an extra component in the cosmological fluid - Dark energy  $\rightarrow$   $\land$ CDM model

ightarrow The theory of gravitation on large scales is not GR

New "modified gravity" theory not found yet

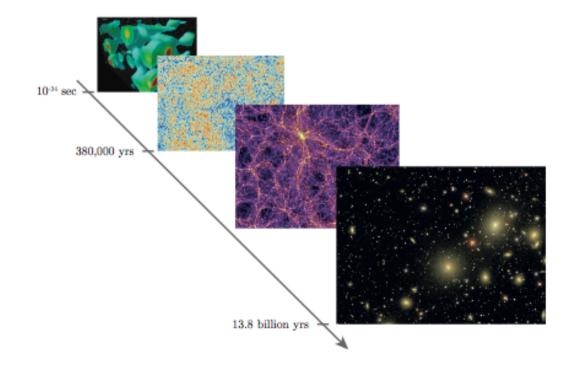


Awarded to Saul Perlmutter, Brian P. Schmidt and Adam G. Riess "for the discovery of the accelerating expansion of the Universe through observations of distant supernovae." (1/2 + 1/4 + 1/4)



**In Summary:** General Relativity + Big Bang + Inflation + Gravitational clustering + cosmological fluid that includes dark matter of the type cold and dark energy of the type cosmological constant.

This physical model has been the standard model of the Universe since the beginning of the XXI<sup>st</sup> century and it is known as  $\Lambda$ CDM.

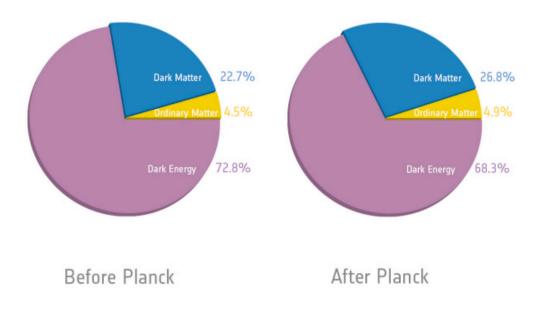


 $\Lambda$ CDM is a complex model

It is a theoretical construction supported by observations.

It includes a variety of physical processes that occur in a variety of epochs, in a variety of scales and contains a large number of free parameters.

The values of the cosmological parameters determine the details of the expansion of the Universe and the evolution and formation of its large-scale structures  $\rightarrow$  they determine the "cosmology".



	Planck+WP		Planck+WP+highL		Planck+lensing+WP+highL		Planck+WP+highL+BAO	
Parameter	Best fit	68% limits	Best fit	68% limits	Best fit	68% limits	Best fit	68% limits
$\Omega_{\rm b} h^2$	0.022032	$0.02205 \pm 0.00028$	0.022069	$0.02207 \pm 0.00027$	0.022199	$0.02218 \pm 0.00026$	0.022161	0.02214 ± 0.00024
$\Omega_c h^2$	0.12038	$0.1199 \pm 0.0027$	0.12025	$0.1198 \pm 0.0026$	0.11847	$0.1186 \pm 0.0022$	0.11889	$0.1187 \pm 0.0017$
1009 <sub>MC</sub>	1.04119	$1.04131 \pm 0.00063$	1.04130	$1.04132 \pm 0.00063$	1.04146	$1.04144 \pm 0.00061$	1.04148	$1.04147 \pm 0.00056$
τ	0.0925	$0.089^{+0.012}_{-0.014}$	0.0927	0.091+0.013 -0.014	0.0943	0.090+0.013 -0.014	0.0952	$0.092 \pm 0.013$
<i>n</i> <sub>s</sub>	0.9619	$0.9603 \pm 0.0073$	0.9582	$0.9585 \pm 0.0070$	0.9624	$0.9614 \pm 0.0063$	0.9611	$0.9608 \pm 0.0054$
$\ln(10^{10}A_s)$	3.0980	3.089 <sup>+0.024</sup> <sub>-0.027</sub>	3.0959	$3.090 \pm 0.025$	3.0947	$3.087 \pm 0.024$	3.0973	$3.091 \pm 0.025$
A <sup>PS</sup> <sub>100</sub>	152	$171 \pm 60$	209	$212 \pm 50$	204	$213 \pm 50$	204	$212 \pm 50$
A <sup>PS</sup> <sub>143</sub>	63.3	$54 \pm 10$	72.6	73 ± 8	72.2	72 ± 8	71.8	$72.4 \pm 8.0$
A <sup>PS</sup> <sub>217</sub>	117.0	$107^{+20}_{-10}$	59.5	59 ± 10	60.2	58 ± 10	59.4	59 ± 10
A <sup>CIB</sup> <sub>143</sub>	0.0	< 10.7	3.57	$3.24 \pm 0.83$	3.25	$3.24 \pm 0.83$	3.30	$3.25 \pm 0.83$
A <sup>CIB</sup> <sub>217</sub>	27.2	29+6	53.9	$49.6 \pm 5.0$	52.3	$50.0 \pm 4.9$	53.0	$49.7 \pm 5.0$
A <sup>tSZ</sup> <sub>143</sub>	6.80		5.17	$2.54^{+1.1}_{-1.9}$	4.64	$2.51^{+1.2}_{-1.8}$	4.86	2.54+1.2
r <sup>PS</sup> <sub>143×217</sub>	0.916	> 0.850	0.825	0.823+0.069	0.814	$0.825 \pm 0.071$	0.824	$0.823 \pm 0.070$
r <sup>CIB</sup> <sub>143×217</sub>	0.406	$0.42 \pm 0.22$	1.0000	> 0.930	1.0000	> 0.928	1.0000	> 0.930
γ <sup>CIB</sup>	0.601	0.53+0.13 -0.12	0.674	$0.638 \pm 0.081$	0.656	$0.643 \pm 0.080$	0.667	$0.639 \pm 0.081$
ξ <sup>tSZ×CIB</sup>	0.03		0.000	< 0.409	0.000	< 0.389	0.000	< 0.410
A <sup>kSZ</sup>	0.9		0.89	5.34+2.8	1.14	4.74+2.6	1.58	5.34+2.8
Ω <sub>Λ</sub>	0.6817	0.685+0.018	0.6830	0.685+0.017	0.6939	0.693 ± 0.013	0.6914	0.692 ± 0.010
σ <sub>8</sub>	0.8347	$0.829 \pm 0.012$	0.8322	$0.828 \pm 0.012$	0.8271	0.8233 ± 0.0097	0.8288	$0.826 \pm 0.012$
Zre	11.37	$11.1 \pm 1.1$	11.38	$11.1 \pm 1.1$	11.42	$11.1 \pm 1.1$	11.52	$11.3 \pm 1.1$
$H_0$	67.04	$67.3 \pm 1.2$	67.15	$67.3 \pm 1.2$	67.94	$67.9 \pm 1.0$	67.77	$67.80 \pm 0.77$
Age/Gyr	13.8242	$13.817 \pm 0.048$	13.8170	$13.813 \pm 0.047$	13.7914	$13.794 \pm 0.044$	13.7965	$13.798 \pm 0.037$
1009	1.04136	$1.04147 \pm 0.00062$	1.04146	$1.04148 \pm 0.00062$	1.04161	$1.04159 \pm 0.00060$	1.04163	$1.04162 \pm 0.00056$
<i>r</i> <sub>drag</sub>	147.36	$147.49\pm0.59$	147.35	$147.47 \pm 0.59$	147.68	$147.67\pm0.50$	147.611	$147.68 \pm 0.45$

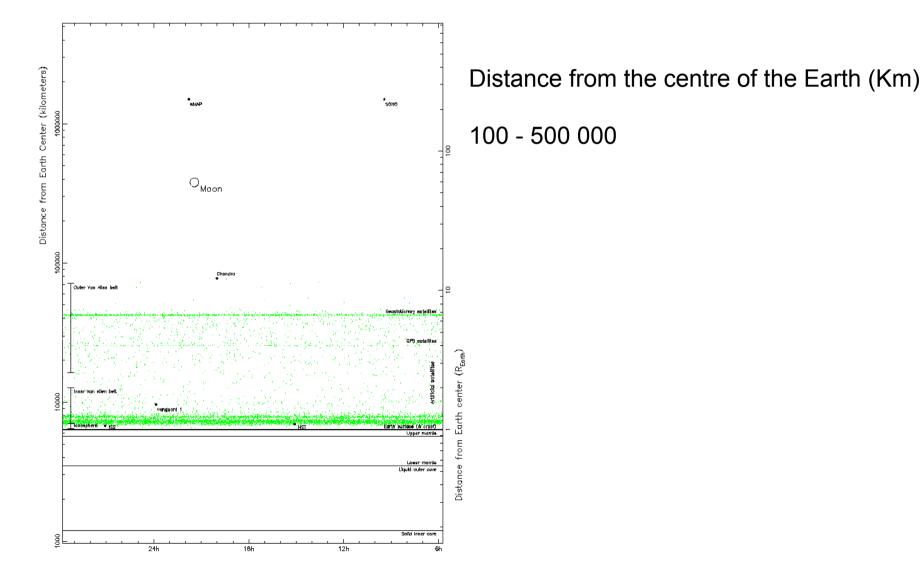
### fundamental cosmological parameters

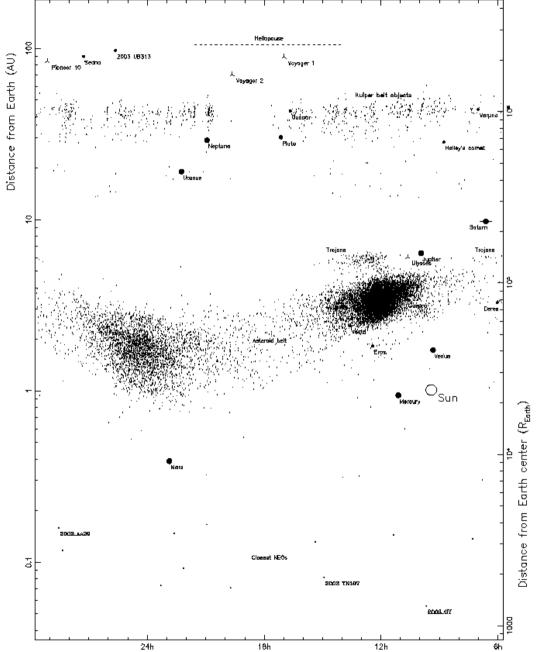
nuisance parameters (of the particular cosmological probe)

### derived cosmological parameters

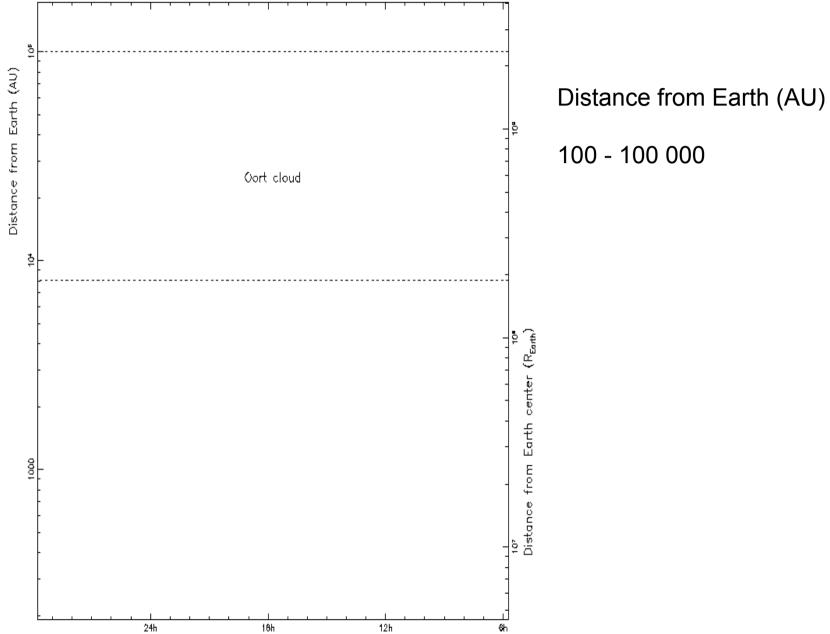
Map of the observed Universe - logarithmic scale and showing the astrophysical objects in their actual coordinates

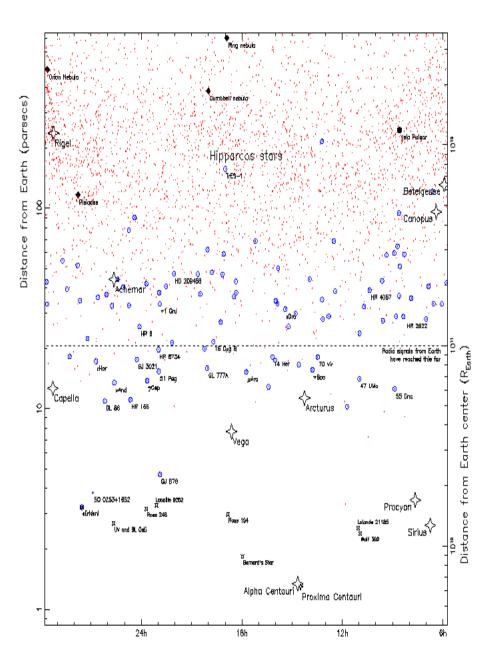
(Gott et al. 2005) http://www.astro.princeton.edu/universe/





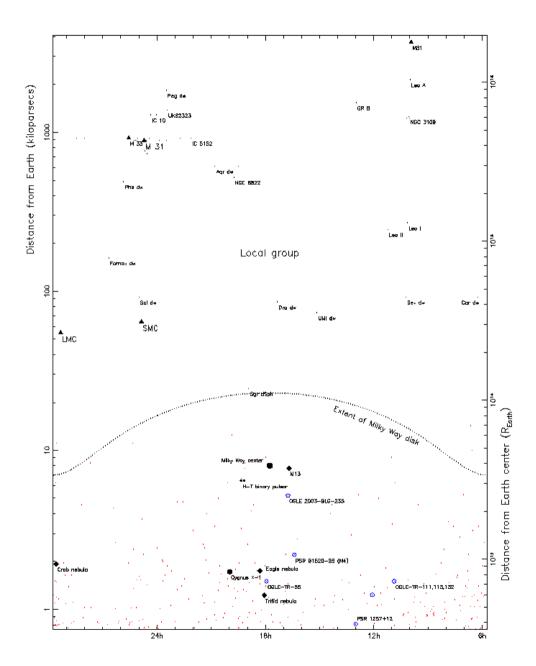
Distance from Earth (AU) 0.05 - 100 (1 AU = 150 000 000 Km)





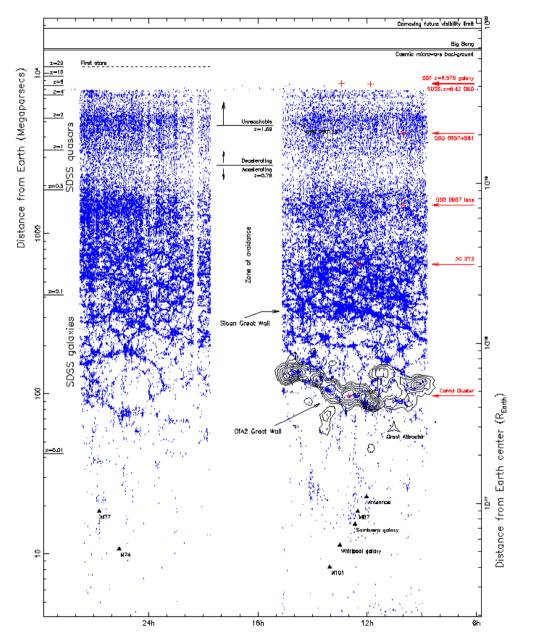
Distance from Earth (pc)

1 - 1000



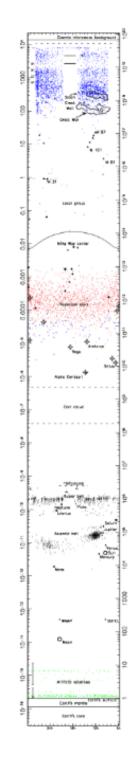
### Distance from Earth (Kpc)

1 - 3000



Distance from Earth (Mpc)

5 - 10 000





Artistic image of Gott's map (P. Budassi 2015)

### Movie of the observed Universe

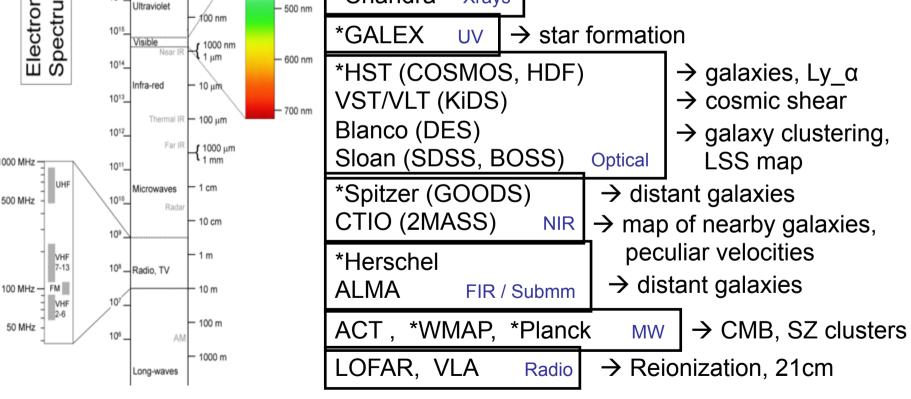
(the Rubin Museum of Art and the American Museum of Natural History) https://www.youtube.com/watch?v=17jymDn0W6U



# The Known Universe

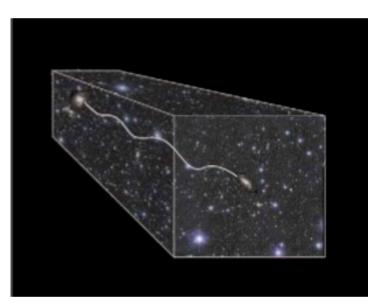
# How are these data obtained and how are they used to find the parameter values?

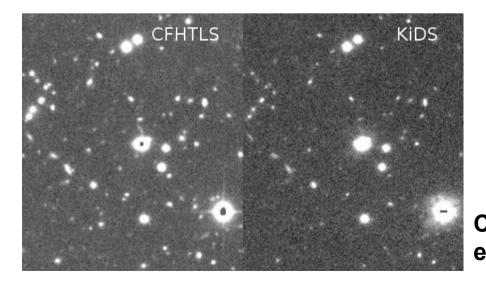
### **Observatories** (ground and \*space) and **Surveys** (FH) Most of the Visible light Long-wavelength Frequency infrared spectrum Radio waves observable observable radio waves Gamma rays, X-rays and ultraviolet absorbed by from Earth from Earth. blocked light blocked by the upper atmosphere atmospheric with some (best observed from space). gasses (best atmospheric observed distortion from space) - 0.1 Å 100 9 Gamma-rays 1019 Atmospheric opacity **1**1Å Electromagnetic 1<sub>0.1 nm</sub> 50 % 1018 X-rays - 1 nm 1017 \*XMM 0 % 1 nm 10 nm 100 nm 1 µm 10 µm 100 µm 1 mm 1 cm 10 cm 10 m 100 m 10 nm Spectrum $\rightarrow$ clusters Wavelength \*Chandra **Xrays** 1016 Ultraviolet 500 nm 100 nm 1015 \*GALEX star formation UV $\rightarrow$ Visible 1000 nm l 1 μm - 600 nm 1014\_ \*HST (COSMOS, HDF) $\rightarrow$ galaxies, Ly $\alpha$ Infra-red 10 µm VST/VLT (KiDS) $\rightarrow$ cosmic shear 1013 700 nm Thermal IR - 100 µm Blanco (DES) 1012\_ Far IR f 1000 µm Sloan (SDSS, BOSS) Optical LSS map ι<sub>1 mm</sub> 1000 MHz 1011\_ UHF \*Spitzer (GOODS) $\rightarrow$ distant galaxies • 1 cm Microwaves 500 MHz 10<sup>10</sup>\_

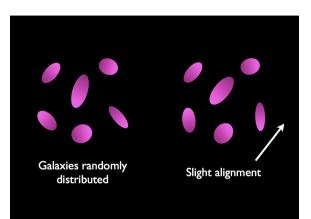


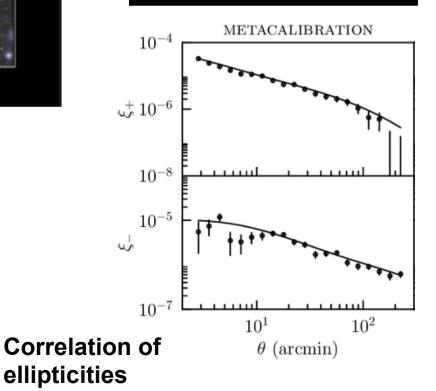
### **Cosmological probes: extracting cosmological information from the data**

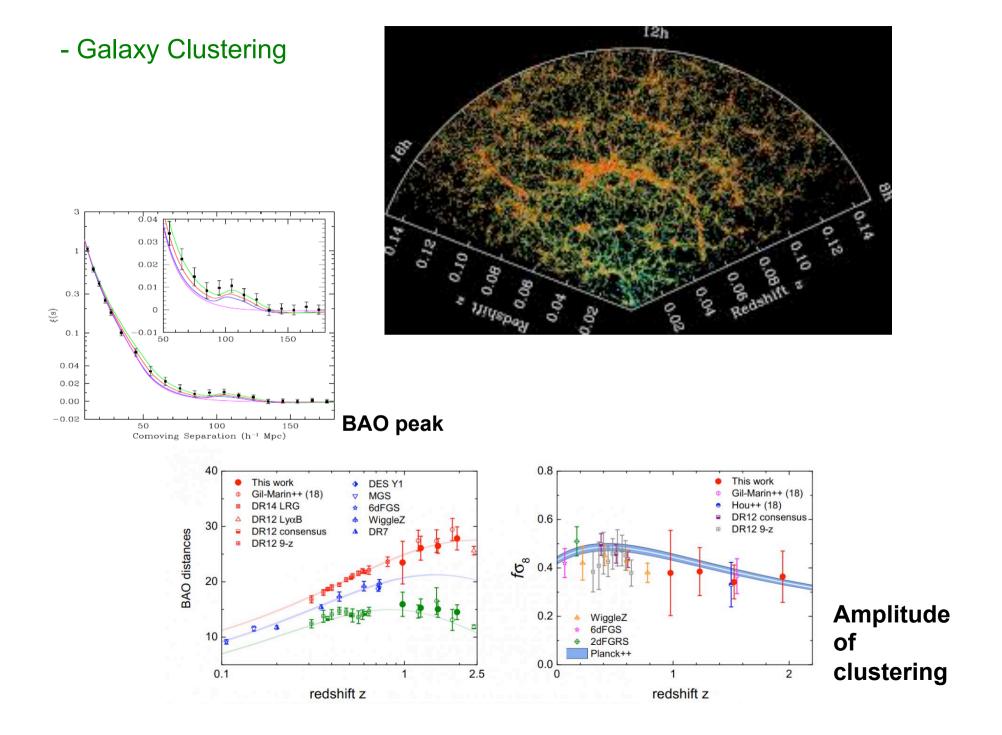
- Weak Lensing

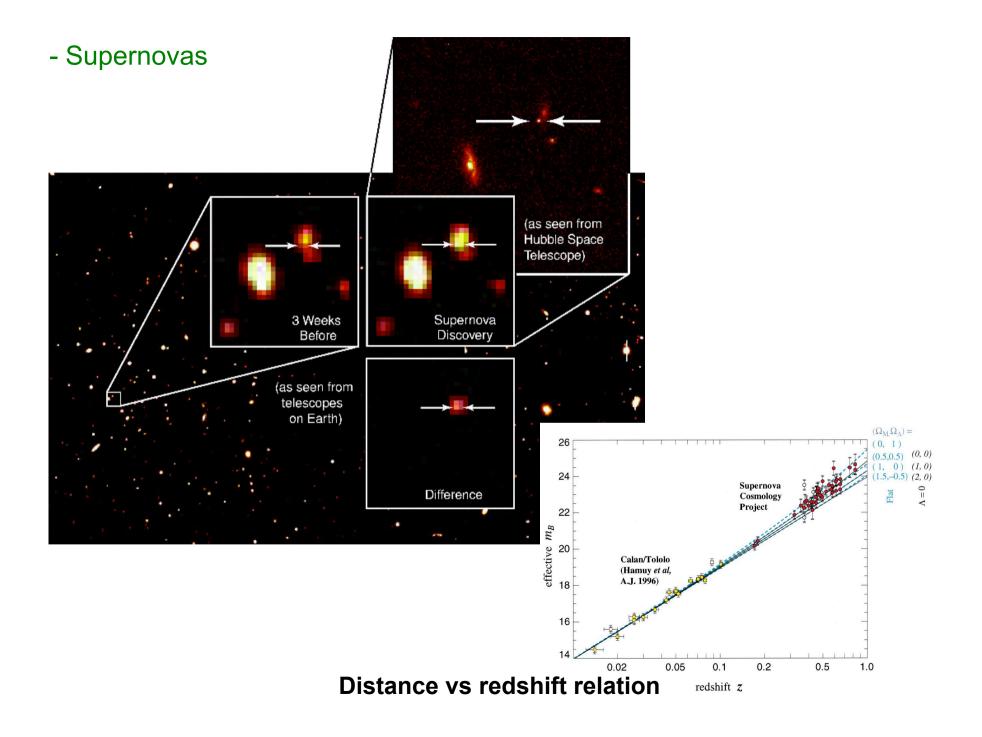




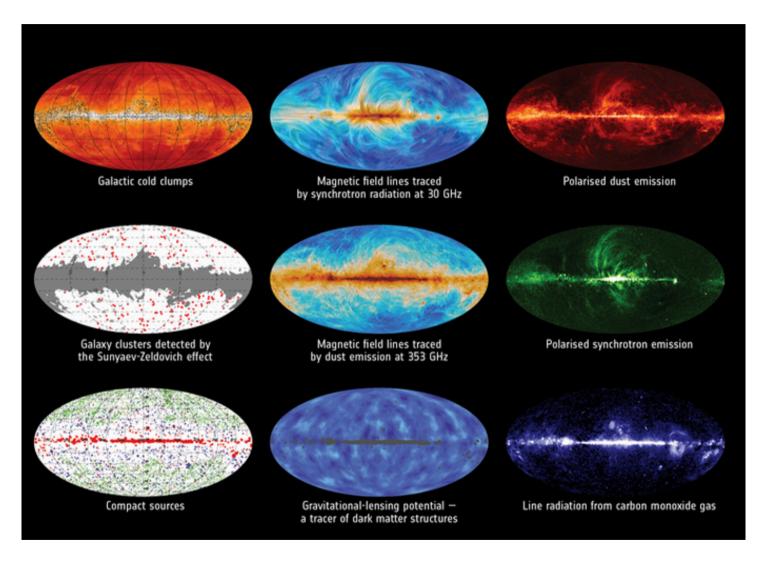




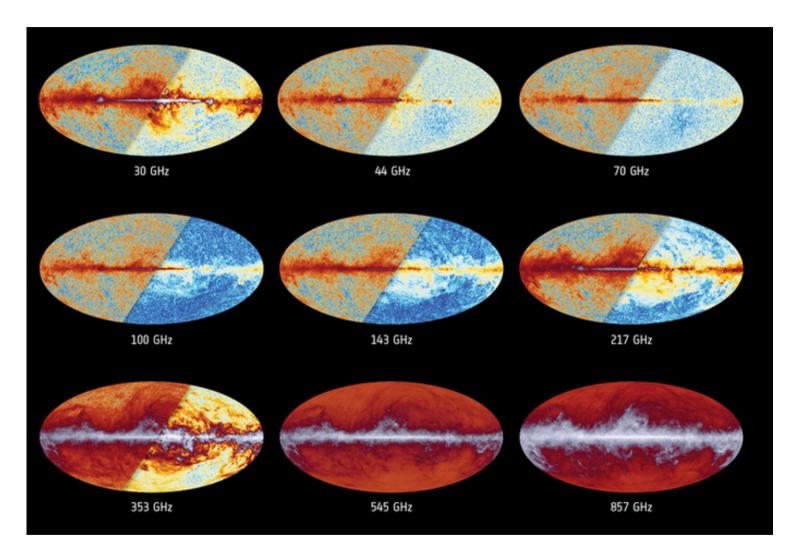




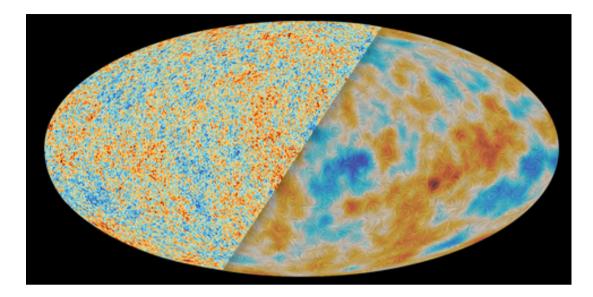
### - Cosmic Microwave Background



(Foregrounds)

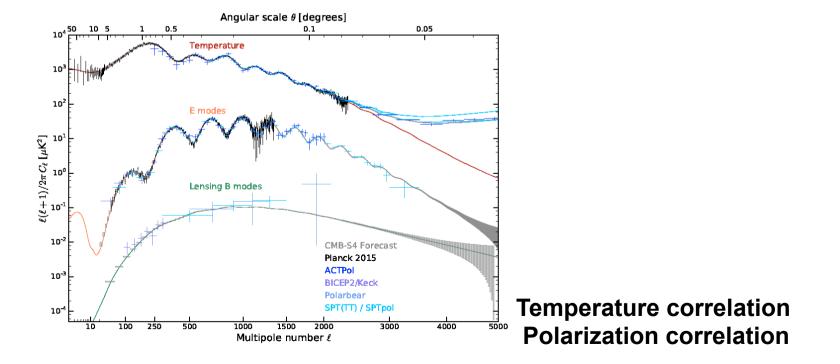


(Cosmological signal in each channel)

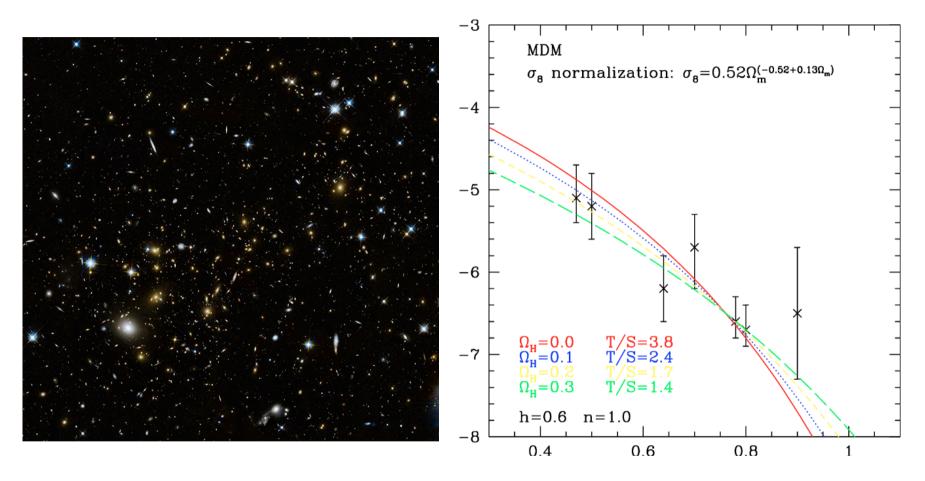


Temperature and polarization maps:

combining 9 frequencies, removing foregrounds, reconstructing the galaxy plane



# - Clusters of galaxies



Cluster abundance (mass function)

and many other cosmological probes:

- Lyman alpha (probes the intergalactic medium)
- 21 cm maps (probes neutral hydrogen)
- Redshift drift (direct measurement of the expansion)
- Time-delays from double images (probes the geometry of space)

- ...

In the near future: new cosmological observations already planned in the European and American programs will obtain *maps of the extra-galactic sky on different bands, at different redshifts, using telescopes of different apertures and field-of-views.* 

# American programs:

National Research Council Decadal Survey 2020

- Vera C. Rubin observatory (2022): imaging (previously named LSST "Large Synoptic Survey Telescope")

- Nancy Grace Roman space telescope (2025): space (previously named WFIRST "Wide-Field Infra-Red Survey Telescope")

Department of Energy / National Science Foundation

- DESI "Dark Energy Spectroscopic Instrument" (2020): spectroscopic

Explorers Program (NASA)

- SPHEREx "Spectro-Photometer for the History of the Universe, Epoch of Reionization, and Ices Explorer" (2023): space

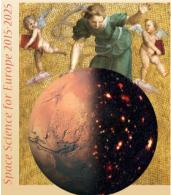




# **European program:** Cosmic Vision (ESA)



# Cosmic Vision



**Euclid Consortium Meeting 2016** Mapping the geometry of the dark universe Centro Cultural de Belém, Lisboa, Portugal 30 May - 03 June



The annual meetings of the Euclid Consortium provide the opportunity to bring together the consortium 1000-plus members and ESA and aerospace industry partners for an update on the mission science, instruments, ground segment, surveys, payload and spacecraft activities



8 space missions

2020 - M1 - Solar Orbiter

2022 - M2 - Euclid (PT: António da Silva, Ismael Tereno)

2022 - L1 - JUICE "JUpiter Icy moons Explorer"

2026 - M3 - PLATO "PLAnetary Transits and Oscillations of stars" (PT: Nuno Santos)

2028 - M4 - ARIEL "Atmospheric Remote-sensing Exoplanet Large-survey" (PT: Pedro Machado)

2031 - L2 - ATHENA "Advanced Telescope for High-ENergy Astrophysics"

2032 - M5 - TBD among EnVision, THESEUS "Transient High Energy Sky and Early Universe Surveyor", SPICA "SPace Infrared telescope for Cosmology and Astrophysics"

2034 - L3 - LISA "Laser Interferometer Space Antenna"

Nowadays there is much interaction between theory and observations.

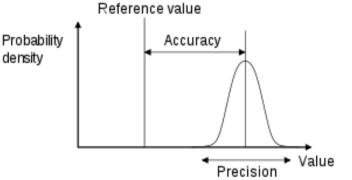
Modern research in cosmology deals with:

### - Testing cosmological models

Compute theoretical predictions of  $\Lambda\text{CDM}$  or alternative models for various observables

Find an explanation for the tensions that exist between various observations (such as different values of  $H_0$  when obtained from local measurements or from CMB observations; different value of  $\sigma$ 8 from CMB and LSS measurements). Does it come from experimental systematic effects of is it an evidence for the need for a new model (or new physics) ?

Determine the values of the cosmological parameters with increasing precision and accuracy



### - Make theoretical improvements to ΛCDM

Build new models of dark energy

Characterize the relation between dark and baryonic matter, introducing more baryonic effects in the description of structure formation (interface cosmology-astrophysics)

### - Explain new observed effects

The acceleration of the expansion is caused by dark energy or can it be predicted by a new gravitational theory of large scales (modifications to General Relativity)?

The small-scale anomalies (cusp/core problem and lack of galaxy satellites) are an evidence against cold dark matter?

The CMB anomalies (multipoles alignment, cold spot) are systematic effects or fundamental physics?

The Lithium problem (nucleosynthesis predicts more primordial Lithium than it is observed)

- Caracterize systematic effects that contaminate the observations

Intrinsic alignments of galaxies affect gravitational lensing

Unknown SN absolute magnitudes affect SN distance measurements

### - Define and derive the properties of new observables and estimators

Intensity mapping (21 cm)

Signal-to-Noise properties of different estimators of the same observable

Estimate the gravitational lensing signal from measured ellipticities

### Detect new effects predicted by the ACDM model

Primordial gravitational waves

CMB polarization

### - Plan and build new cosmological surveys

Future space missions

### - Understand the nature of the ACDM assumptions

What is dark energy (and does it really exist?)

What is dark matter (and does it really exist?)

Did inflation really happen?

Test the cosmological principle

### - Understand the beginning of the Universe

Problem of the initial singularity

Quantum gravity